Desired ultimate beams for Probing BSM at colliders: Luminosity vs energy

LianTao Wang University of Chicago

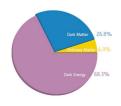
This talk

- Focus on high energy colliders.
- Mainly order of magnitude estimates and scaling.
 - Focus on broad classes of signals, not special models. (There is not a model much more compelling than this others.)
- Both the bare minimal requirement, and what's needed for a more comprehensive program.

Many more detailed studies still needed to be done. I hope to give the impression of the order of magnitude here.

Where we are

- We have a good picture of the universe, including its basic content and the interactions.
 - Standard Model of particle physics
 - Dark matter + dark energy



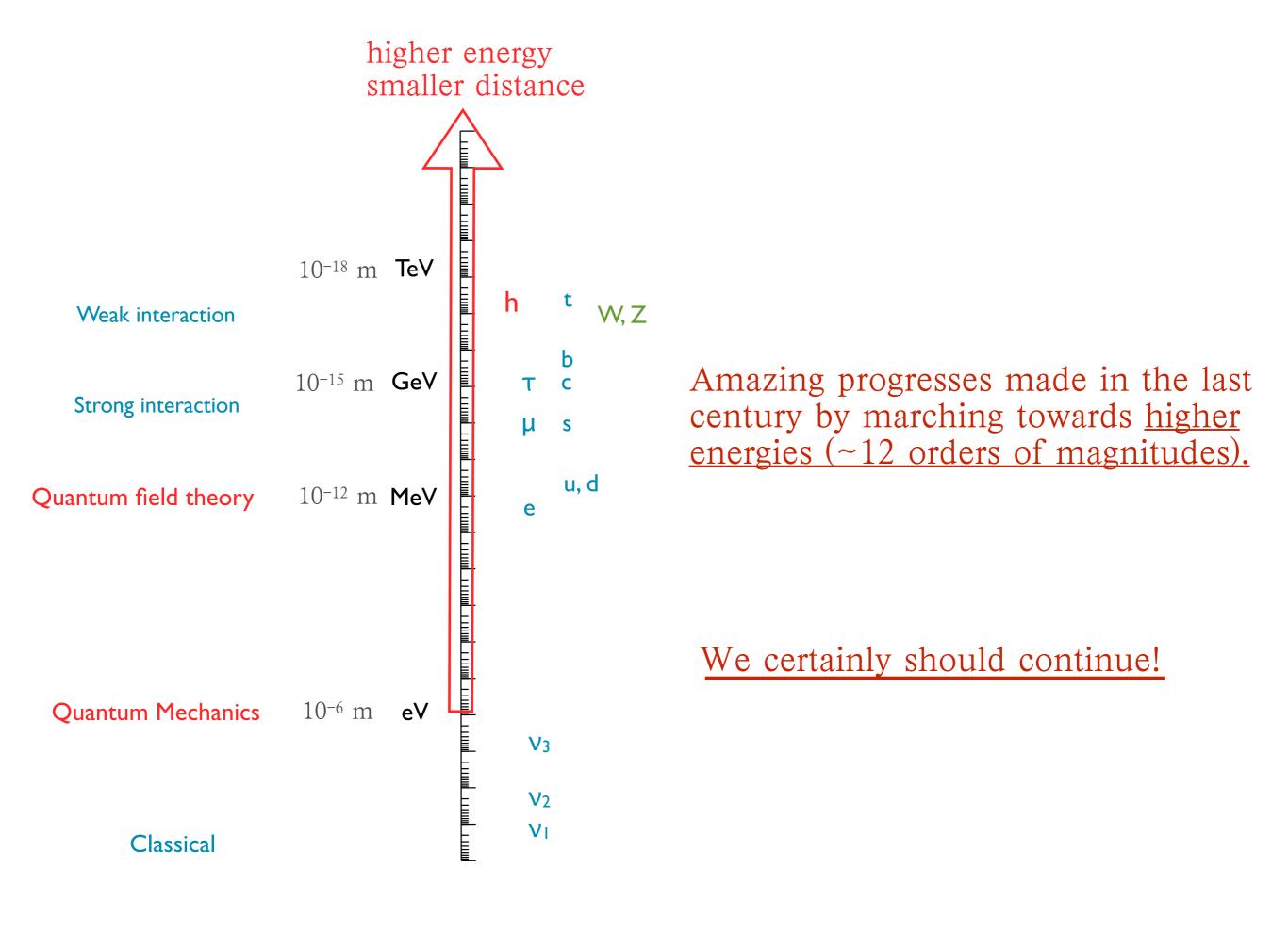
- However, we are missing the underlying principles of this picture.

Many remaining puzzles

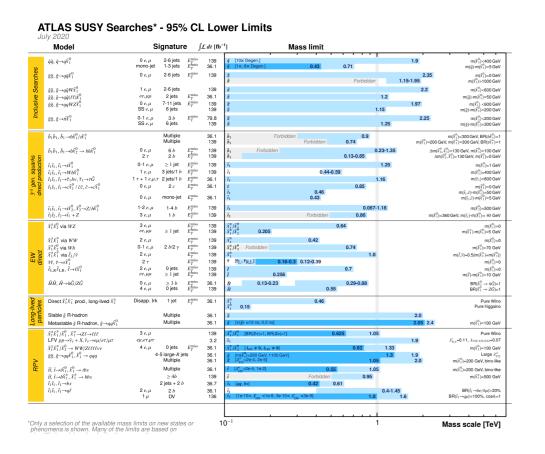
- Origin of the Electroweak scale.
- What is Dark Matter
- What is the origin of Flavor/CP
- Matter anti-matter asymmetry.
- Dark energy?

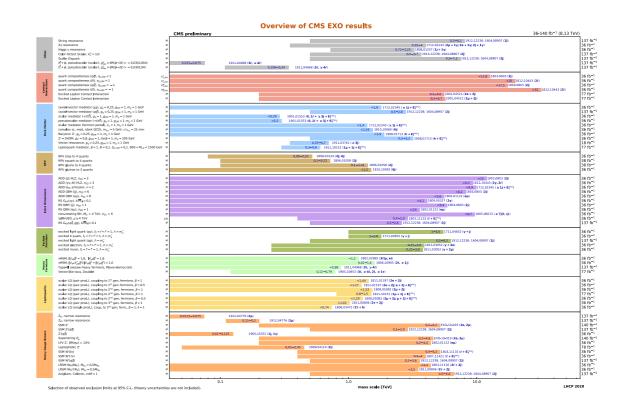
— ...

We believe that going to higher energy (shorter distance) will give us answers.

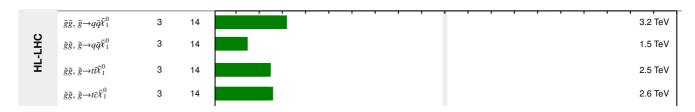


Current status and next decade







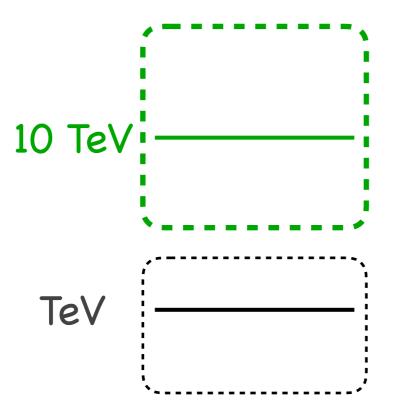


- TeVs explored, coverage will improve at the HL-LHC (by about 50-100%).

The next frontier



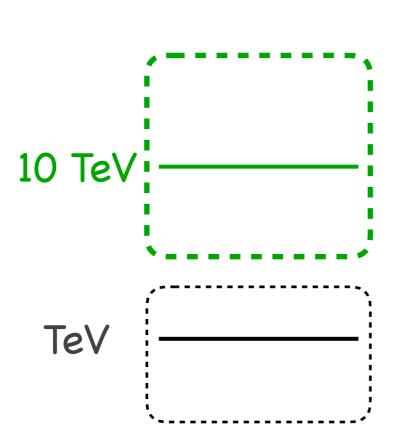
- Naturalness



For decades, we hope naturalness should give us new physics below or around TeV scale.

Could still be there, just harder to discover. Gaps/difficult spots in searches: twin Higgs, ...

- Naturalness



Plausible scenarios have emerged for new physics to be here.
mini-split SUSY, meso-tuned, ...

Also possible that first hint at TeV, but more interesting dynamics here.

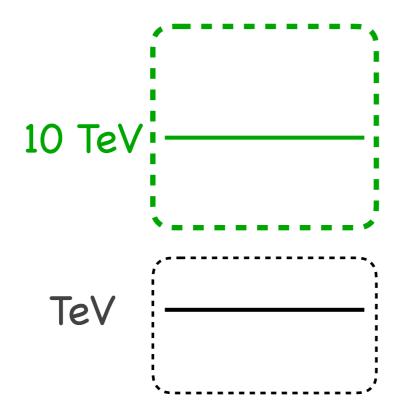
Compositeness, ...

Higher rates, helps cover difficult spots in TeV searches.

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- Dark Matter

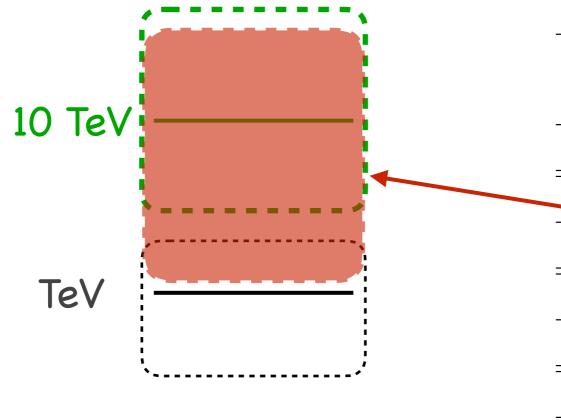


Simplest WIMP dark matter models: dark matter part of a EW multiplet

$egin{array}{c} \operatorname{Model} \ (\operatorname{color}, n, Y) \end{array}$		Therm. target
(1,2,1/2)	Dirac	1.1 TeV
(1,3,0)	Majorana	2.8 TeV
$\boxed{(1,3,\epsilon)}$	Dirac	2.0 TeV
(1,5,0)	Majorana	14 TeV
$(1,5,\epsilon)$	Dirac	6.6 TeV
(1,7,0)	Majorana	23 TeV
$(1,7,\epsilon)$	Dirac	16 TeV

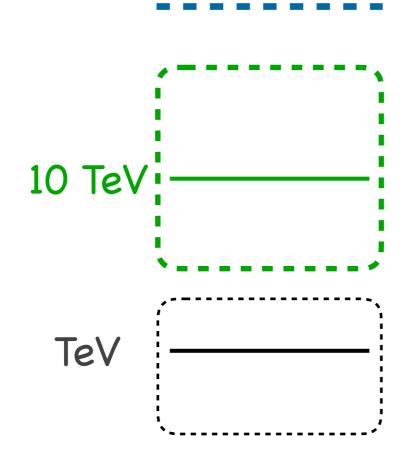
- Dark Matter

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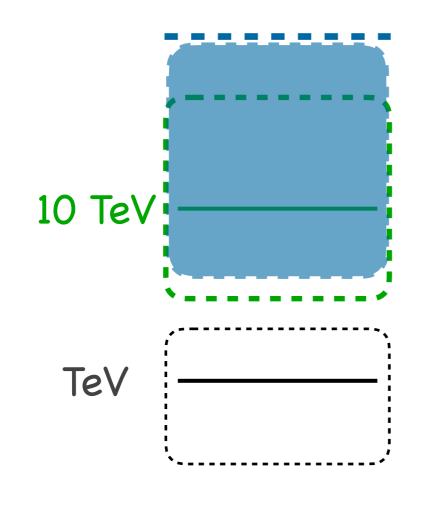
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- Flavor, CP.



Without symmetries, limits already at > 100s TeV

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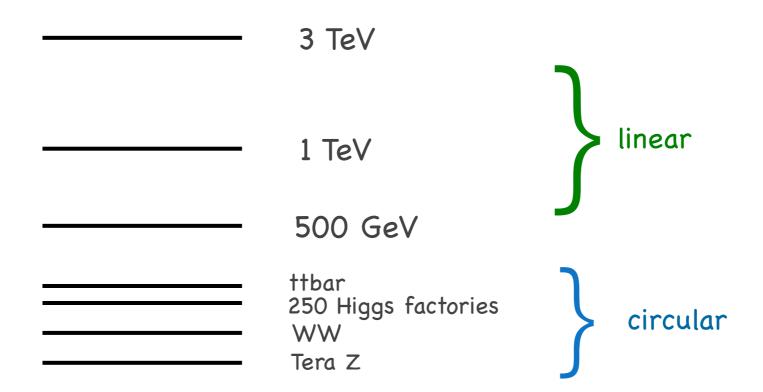
Many scenarios with new flavor physics at 10s TeV. Partial compositeness, some flavor symmetries...

How to get there

- Two possible routes
 - Lepton: e+e-, μ+ μ-
 - \Box $\gamma\gamma$ similar, but somewhat narrower physics program.
 - Hadron, pp.
- Two approaches
 - Direct production of new physics particles.
 - Precision measurement. Can be sensitive to even higher energy scales.

Lepton collider e+e-, µ+µ-

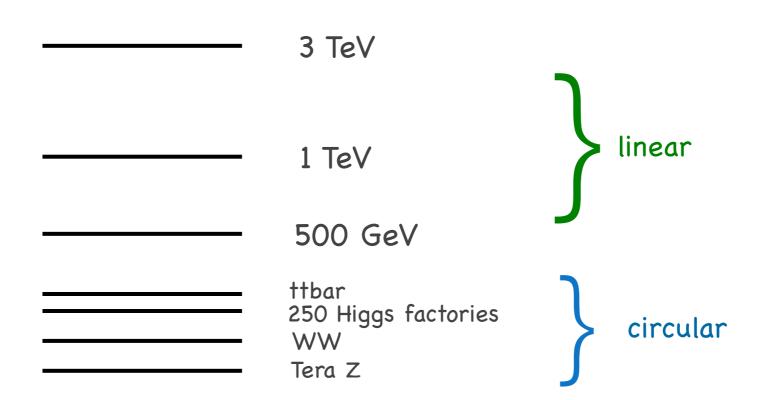
Recent proposals



Physics cases have been studied.

Recent proposals

Beyond this, higher energies? muon collider? $\gamma\gamma$?... What is needed?



Physics cases have been studied.

Direct production

For letpon colliders: $M_{\rm NP} \sim E_{\rm CM}$

in principle no parton distribution suppression, can produce new physics with mass up to beam energy.

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Signal = $\mathscr{L} \cdot \sigma$, \mathscr{L} : luminosity Fixing signal strength: $\to \mathscr{L} \propto E_{\mathrm{CM}}^2$

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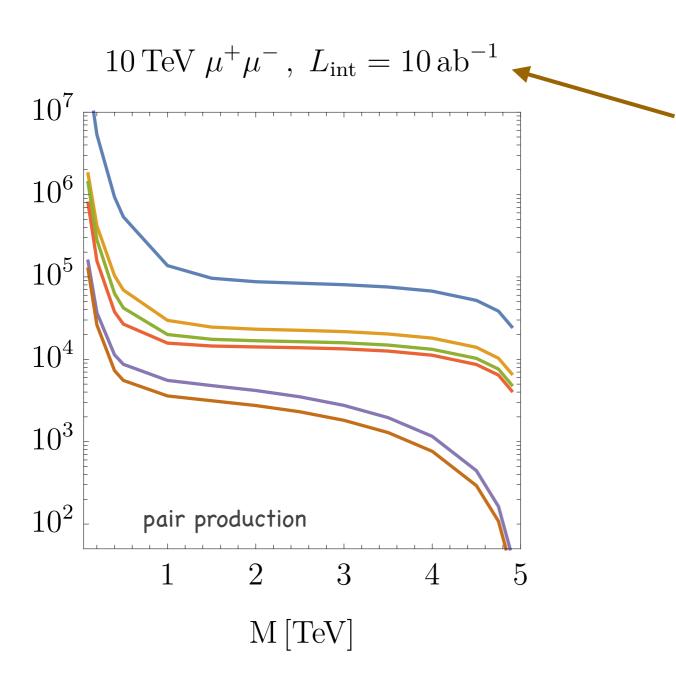
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For higher energies, maybe we just take LEP luminosity and scale up?

More careful consideration

- Discoveries are simpler in general at lepton colliders.
 - New physics give energetic final states.
 - Low background (comparable to signal).
 - ▶ May need only O(10s) signal events.
 - □ Different from the need to produce thousands and millions particles (such as Z boson at LEP) to study them in detail.
- Of course, there are also difficult scenarios to cover.

Top partner search.

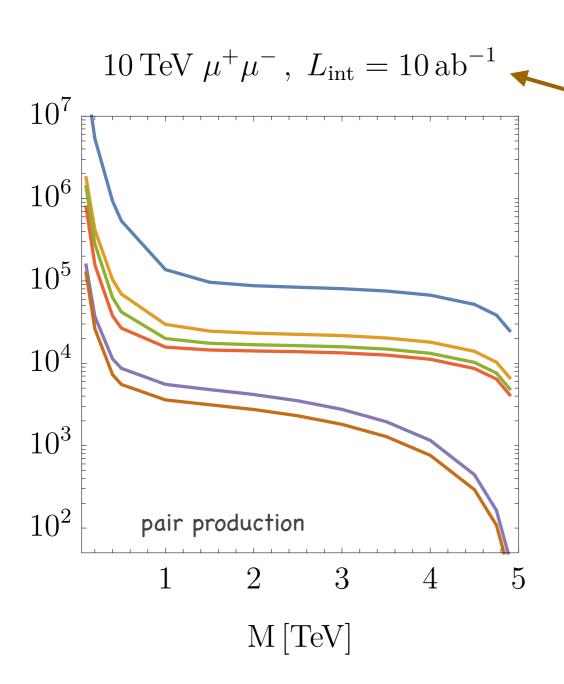


Good production rate up to : $2M \sim E_{\rm CM}$

A benchmark for muon collider

$$\mathcal{L} = \left(\frac{\sqrt{s}}{10 \,\text{TeV}}\right)^2 \times 10^{35} \,\text{cm}^{-2} \text{s}^{-1}$$

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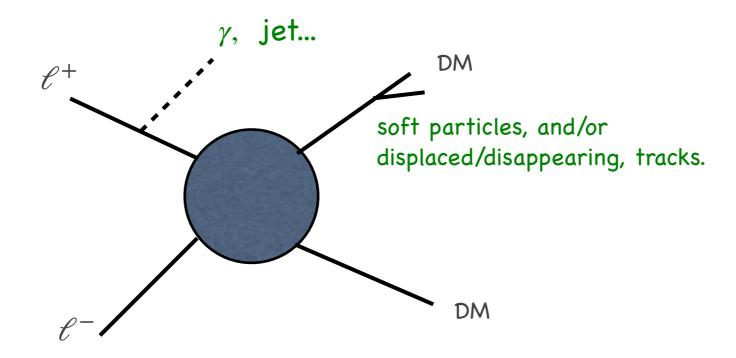
Can have > 105 events

However, for 5 TeV T' with

$$T' \rightarrow tZ, th, bW$$

10s signal events (L=10³² cm⁻² s⁻¹) should be enough.

Harder case, dark matter

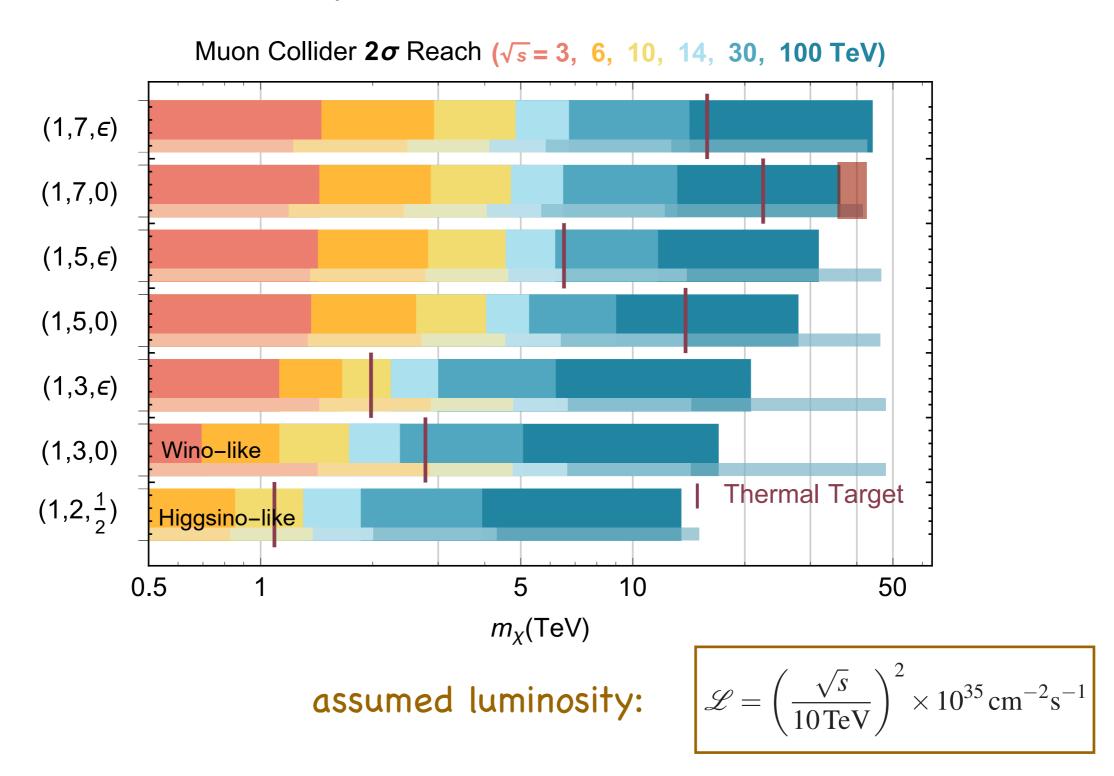


Most energy goes into making the heavy dark matter particle.

The remaining objects, such as photon, jet, soft.

Large background! More difficult.

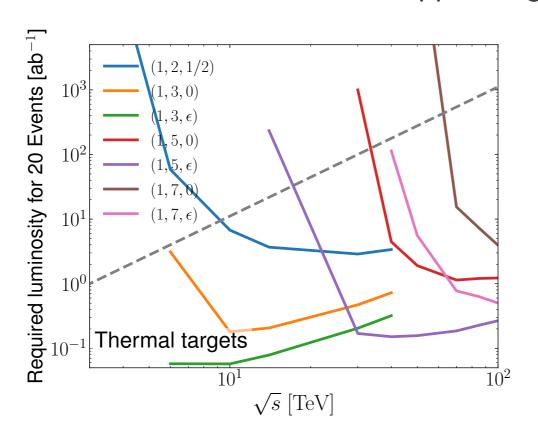
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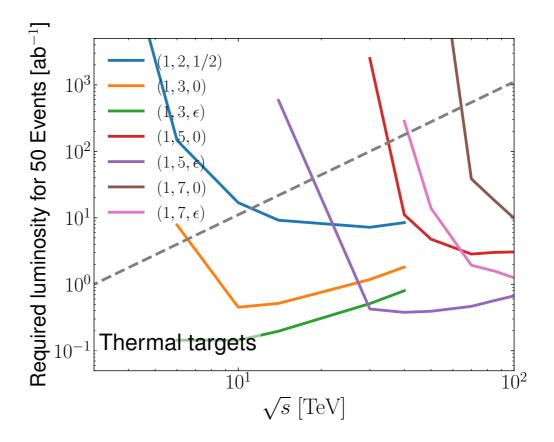


Really need the large luminosity to get there.

Dark matter: E vs L, some trade off

Disappearing track searches





Some examples

doublet: 10 ab-1 at 10 TeV or 3 ab-1 at 20 TeV

Dirac triplet: < 0.1 ab⁻¹ at 6 TeV

Majorana 5-plet: 100 ab⁻¹ at 30 TeV or 1 ab⁻¹ at 100 TeV

Dirac 7-plet: 100 ab⁻¹ at 40 TeV or 10 ab⁻¹ at 50 TeV

Minimal and optimal scenarios

Minimal scenario: can produce at least 10 signal event for weak scale cross section. Can do "basic" new physics searches and cover interesting scenarios.

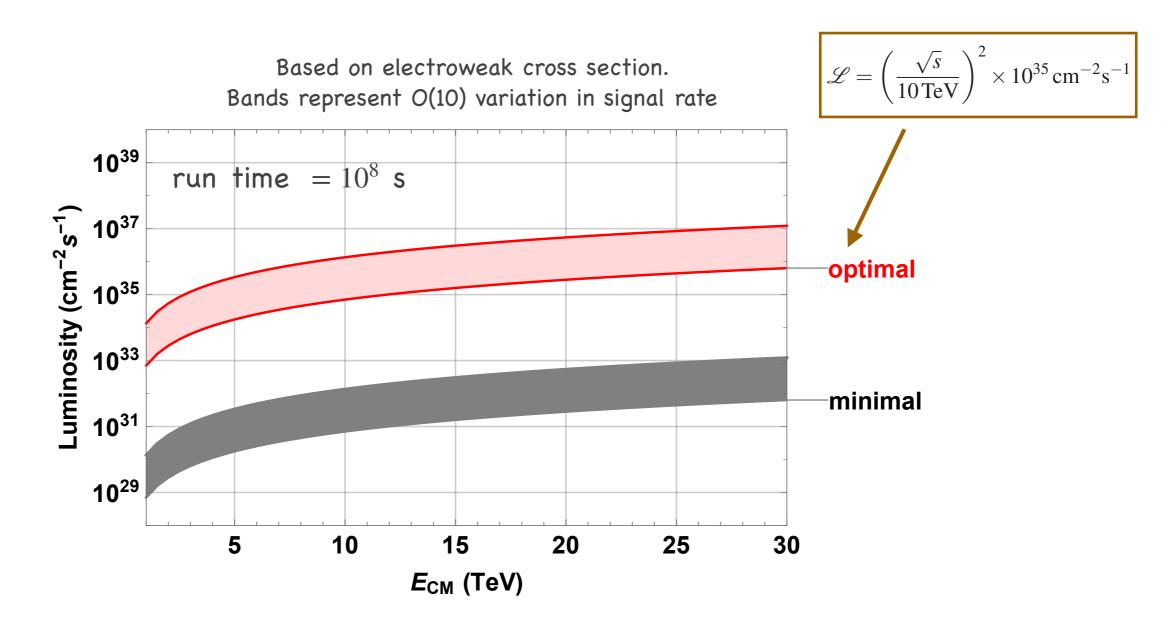
<u>Bare minimum.</u> Will miss some important physics. Maybe only a good starting point.

Optimal scenario: can cover as many difficult cases as possible, such as the dark matter searches.

Some choices needed here, but the basic wishlist is quite commonly accepted.

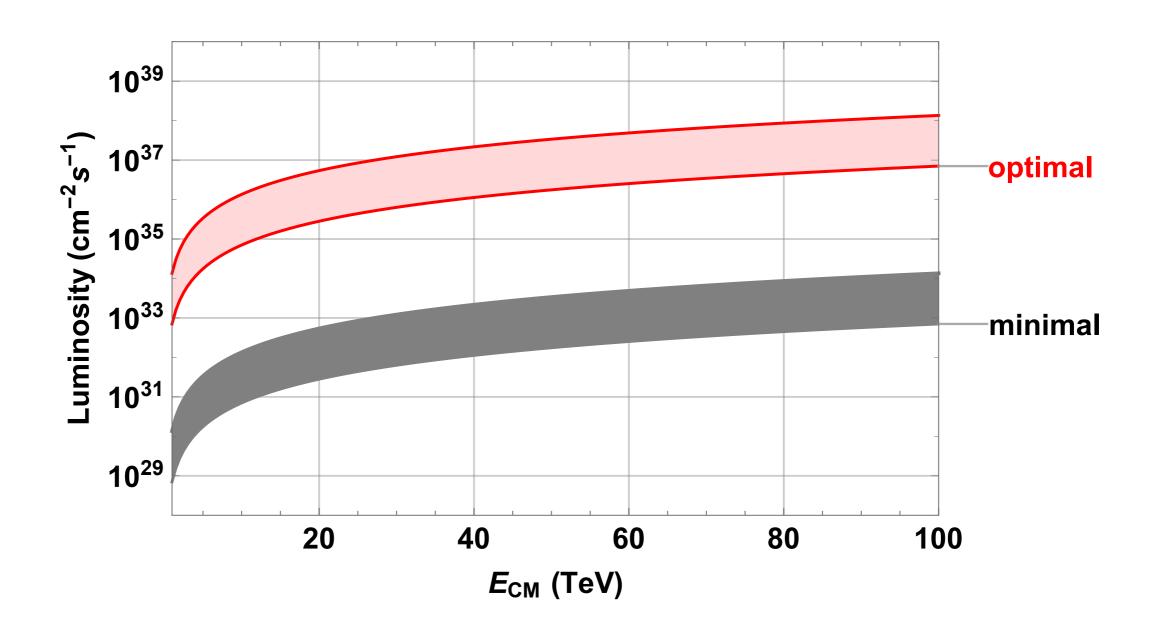
Lepton collider luminosities

- For both muon and electron

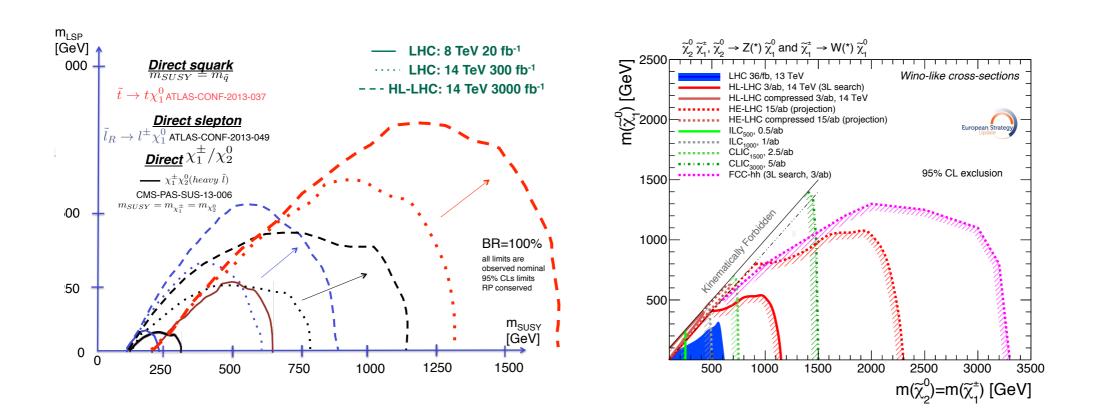


Both scales as $\mathcal{L} \propto E_{\mathrm{CM}}^2$, minimal $\sim 10^{-4} \times \mathrm{optimal}$

Even higher energies



Intermediate energy range 1-10 TeV



For particles without strong interaction, HL-LHC barely covers up to TeV, with gaps in parameter space

Lepton collider can cover up to E_{CM} very well.

Lepton collider below 10 TeV can still improve a lot beyond the LHC.

Precision measurements

- Maybe the new physics scale is above E_{CM} .
- The NP effect can be parameterized by EFT operators

$$\frac{1}{\Lambda^2} \mathcal{O}^{(6)}, \ \frac{1}{\Lambda^4} \mathcal{O}^{(8)}, \dots$$
 $\Lambda \sim \text{scale of new physics}$

- Can only probe through precision measurements.
- Naturally, can be a later part of the physics program.

Precision measurement

Deviation from SM coupling
$$\delta \sim c \frac{m_W^2}{\Lambda^2}, \quad c \sim \mathcal{O}(1)$$

LHC: $\delta \sim a$ few $\% \rightarrow \Lambda \sim \text{TeV}$

 $\delta \sim \mathcal{O}(10^{-3})$ (per mil) needed to reach up to $\Lambda \sim 10$ TeV

Statistics limited:
$$\delta \propto \frac{1}{\mathcal{L}^{1/2}}$$

Higgs coupling measurement needs 106 Higgs at proposed Higgs factories

For heavy new physics parameterized by

$$\frac{1}{\Lambda^2}\mathcal{O}^{(6)}, \quad \frac{1}{\Lambda^4}\mathcal{O}^{(8)}, \dots$$

Effect of new physics larger at higher energy scales

$$(\delta\sigma/\sigma)_{\rm higher~E}\sim \frac{E^2}{\Lambda^2},~~\delta\sigma~{\rm deviation~due~to}~~\mathcal{O}^{(6)}$$

Don't need to do as precise a measurement if we can measure the process at higher energies.

$$(\delta \sigma / \sigma)_{\text{higher E}} \sim \frac{E^2}{\Lambda^2}$$

For example:

At 250 GeV Higgs factories with 106 events

$$(\delta\sigma/\sigma)_{ee} \sim \frac{m_W^2}{\Lambda^2} \sim (10^{-3})_{\text{exp}}$$

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Measuring at energy E, we need

$$\sim \frac{E^2}{m_W^2} \times 10^{-3}$$
 accuracy to reach the same scale

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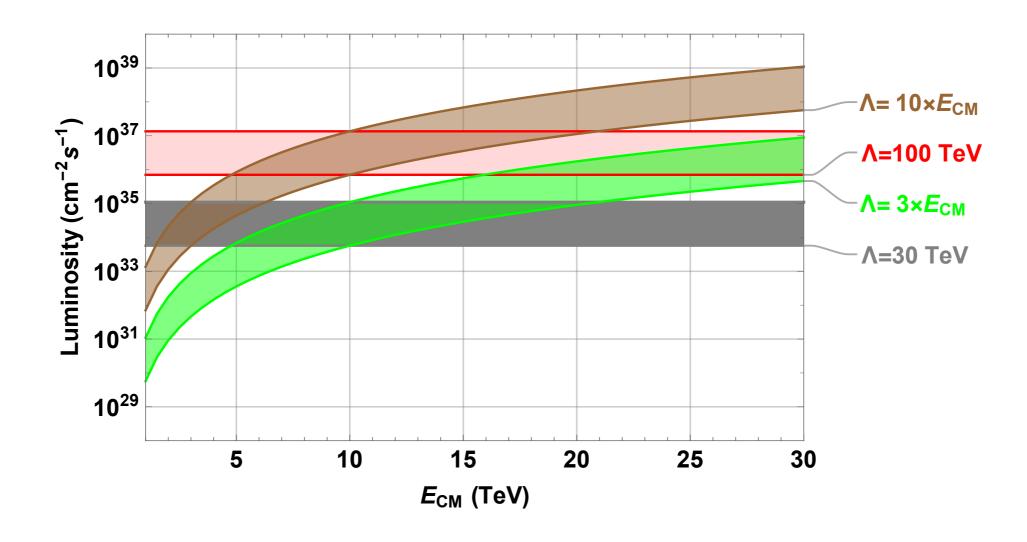
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For example, E=1 TeV, need an accuracy of $10\,\%$ $\rightarrow \sim 100$ events will do.

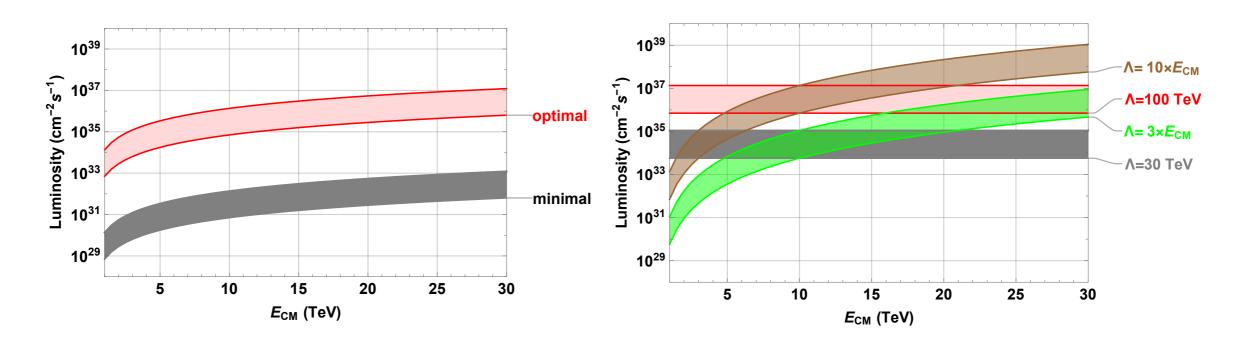
Luminosity need for precision



Minimally, we hope to reach new physics scale $\Lambda \approx 3$ E_{CM}

Optimally, we would like to reach new physics scale $\Lambda \approx 10~\text{E}_{\text{CM}}$ Also cover potential difficult cases.

Lepton collider summary



Luminosity cm ⁻² s ⁻¹	1.5 TeV	3 TeV	6 TeV	10 TeV	I4TeV	30 TeV	100 TeV
Direct search minimal	3×10 ²⁹	1030	5×10 ³⁰	2x10 ³¹	5×10 ³¹	2×10 ³²	2×10 ³³
Direct search optimal	3×10 ³³	1034	5×10 ³⁴	2x10 ³⁵	5×10 ³⁵	2×10 ³⁶	2×10 ³⁷
Precision minimal	3×10 ³⁰	8×10 ³¹	2×10 ³³	1034	5×10 ³⁴	1036	2×10 ³⁷
Precision optimal	7×10 ³²	I 0 ³⁴	2×10 ³⁵	2x10 ³⁶	5x10 ³⁶	1038	2×10 ³⁹

Hadron colliders

LHC and recent proposals

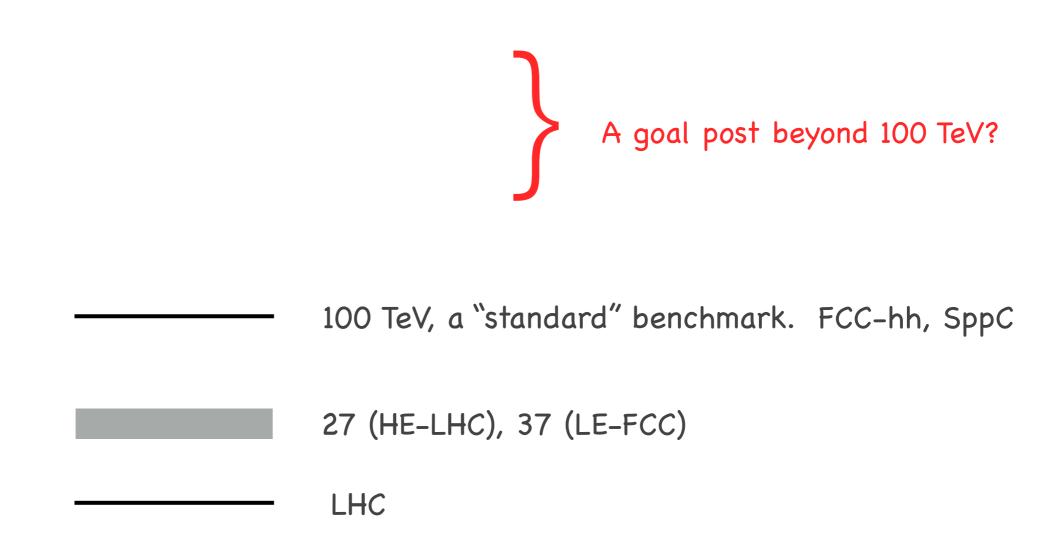
100 TeV, a "standard" benchmark. FCC-hh, SppC

27 (HE-LHC), 37 (LE-FCC)

LHC

Physics case for 27, 37, and 100 TeV have been studied.

LHC and recent proposals



Physics case for 27, 37, and 100 TeV have been studied.

Hadron collider reach

Cross section at hadron collider, for producing heavy new physics with mass M.

$$\sigma \sim L_p \cdot \hat{\sigma} \propto \frac{1}{M^{2a}} \hat{\sigma} \qquad \hat{\sigma} \propto \frac{1}{M^2}$$

Sharp falling Parton Luminosity $L_p \longleftrightarrow a \gg 1$

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For two colliders with different energy and luminosity

$$E_1, \mathcal{L}_1$$
 and E_2, \mathcal{L}_2

Reach in new physics mass, M₁ and M₂ scales as

$$\frac{M_1}{M_2} = \left(\frac{E_1}{E_2}\right)^{\frac{a}{a+1}} \left(\frac{\mathcal{L}_1}{\mathcal{L}_2}\right)^{\frac{1}{2a+2}}$$

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Reach scale with energy, weaker dependence on luminosity

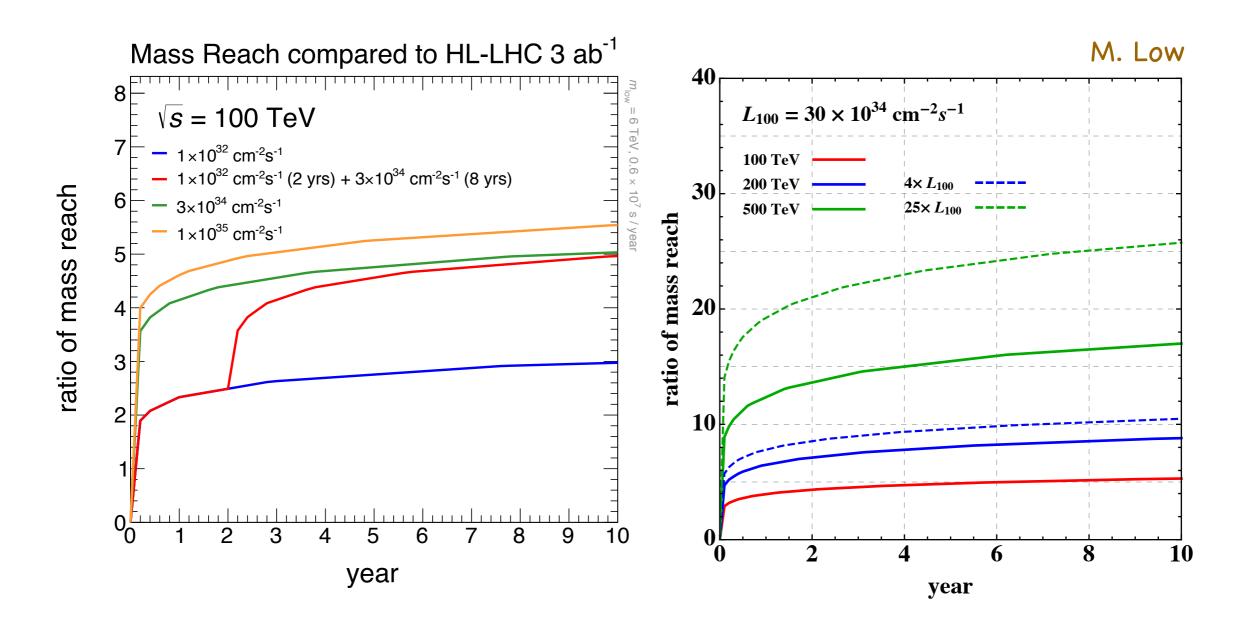
Rapid gain in mass reach after turning on.

Gain slowing down as luminosity increases.

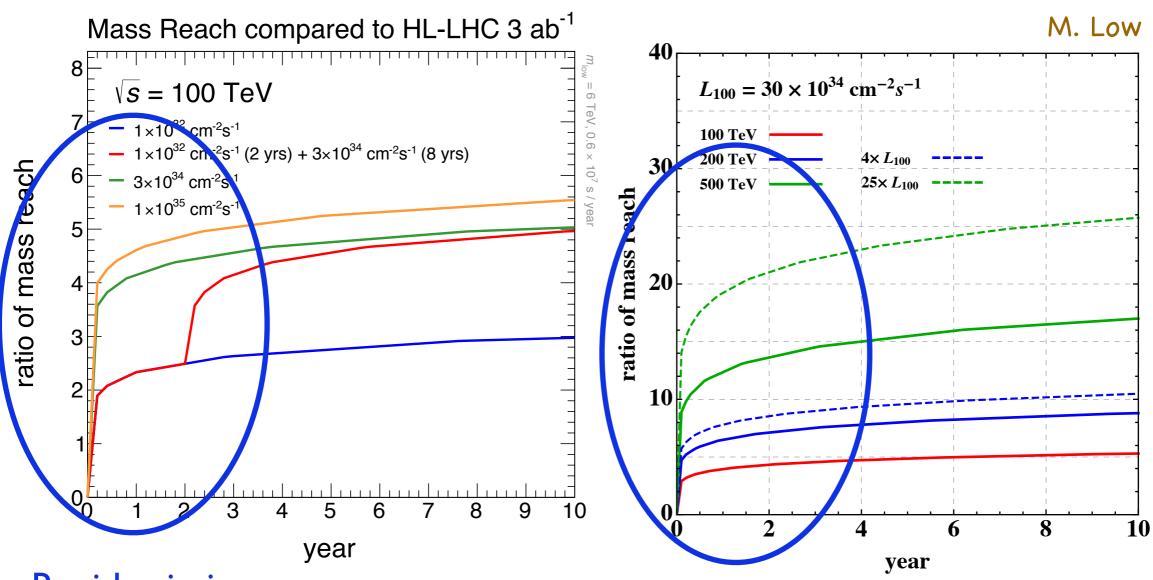
Eventually,
$$\frac{\mathscr{L}_1}{\mathscr{L}_2} = \frac{E_1^2}{E_2^2}, \quad \frac{M_1}{M_2} \to \frac{E_1}{E_2}$$

However, most of the gain come much earlier. Lower luminosity can do a lot already.

Physics program at hadron collider



Physics program at hadron collider

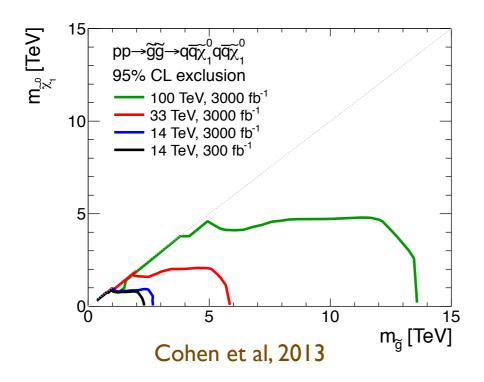


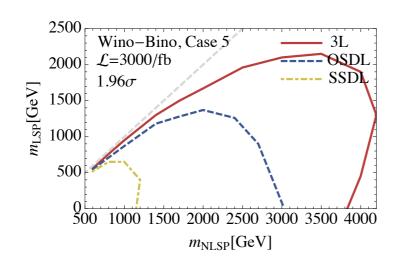
Rapid gain in mass reach

10³⁵-10³⁶ cm⁻²s⁻¹ doing a good job already.

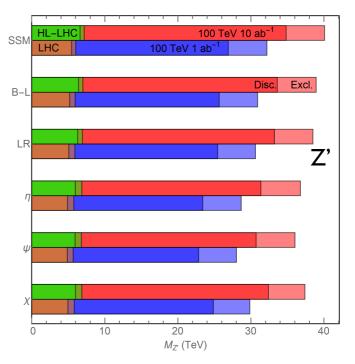
Examples:

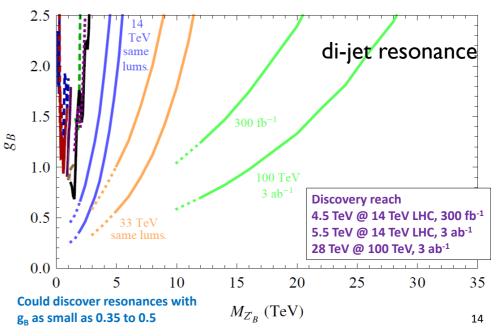
100 TeV: x 5(more) improvement with same lumi as HL-LHC





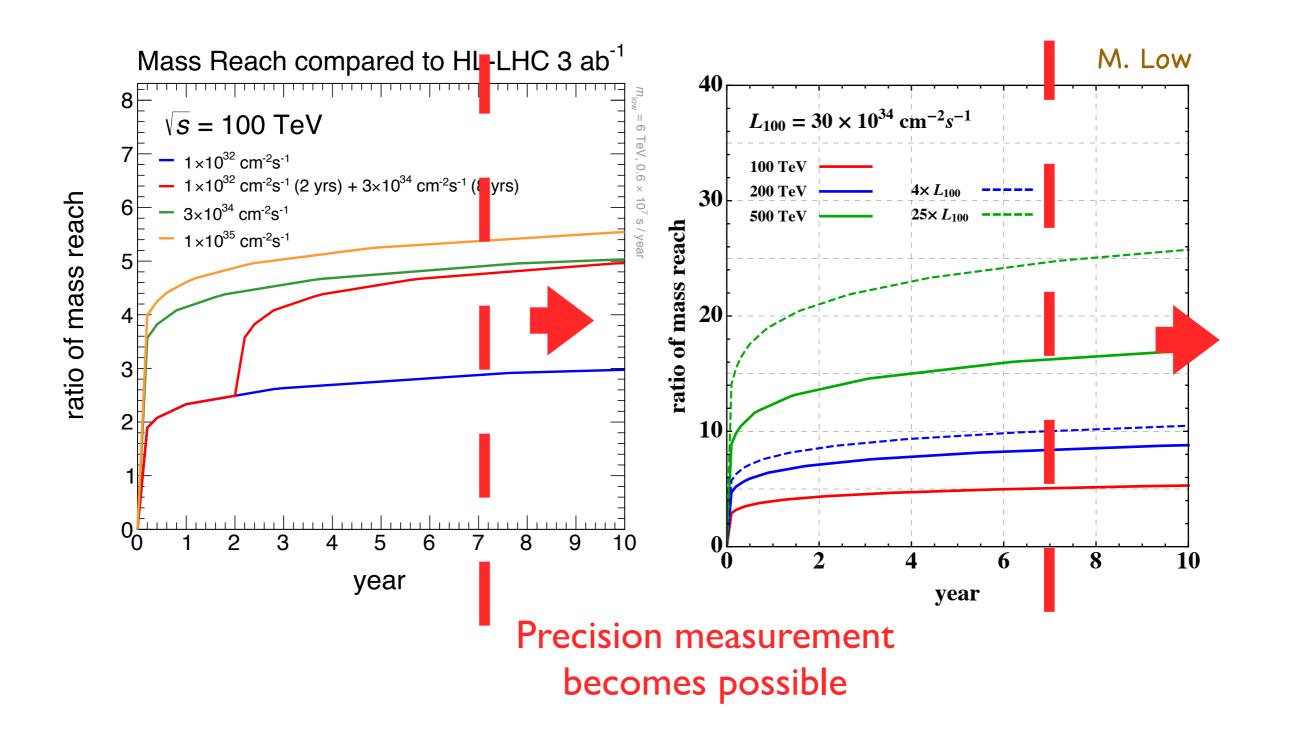
Gori, Jung, LTW, Wells, 2014





Felix Yu, 2013

Physics program at hadron collider



Energy = precision

At FCC-ee/CEPC/ILC

$$(\delta\sigma/\sigma)_{ee} \sim \frac{m_W^2}{\Lambda^2} \sim (10^{-3})_{\text{exp}}$$

At hadron collider

$$(\delta \sigma / \sigma)_{\text{had}} \sim \frac{E^2}{\Lambda^2}$$

Effects larger at higher energies!

 $E = parton energy \approx 0.1 E_{CM}$

Can probe: $\Lambda \sim 0.1 \times E_{\rm CM} \times (\delta \sigma / \sigma)_{\rm exp.error}^{-1/2}$

For example: $(\delta\sigma/\sigma)_{\rm exp.error}\sim 10\,\%$, $\Lambda\sim 30\,$ TeV with $E_{\rm CM}=100\,$ TeV

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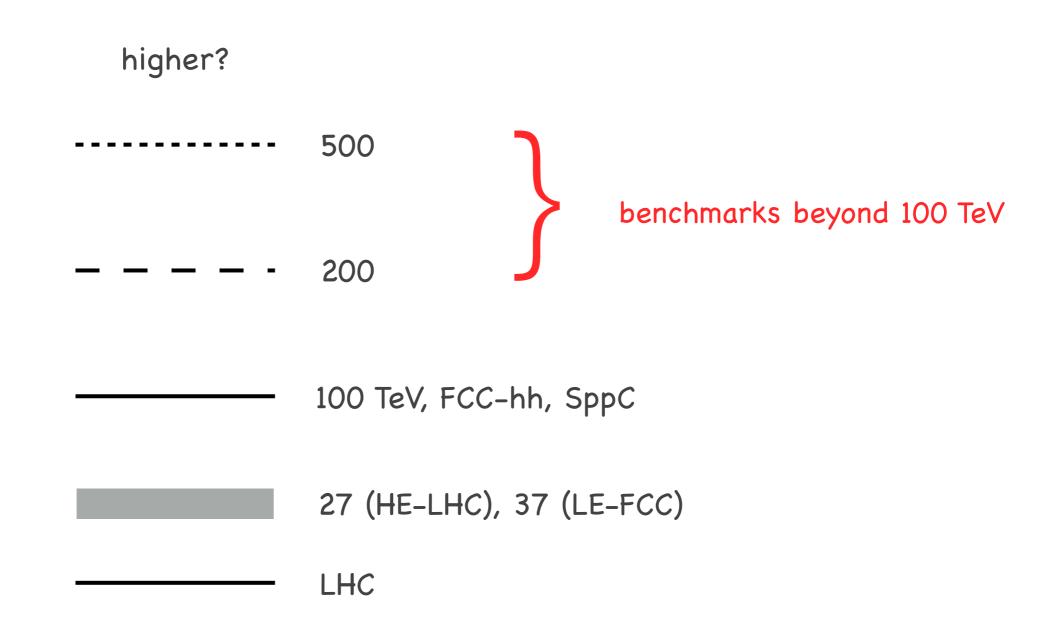
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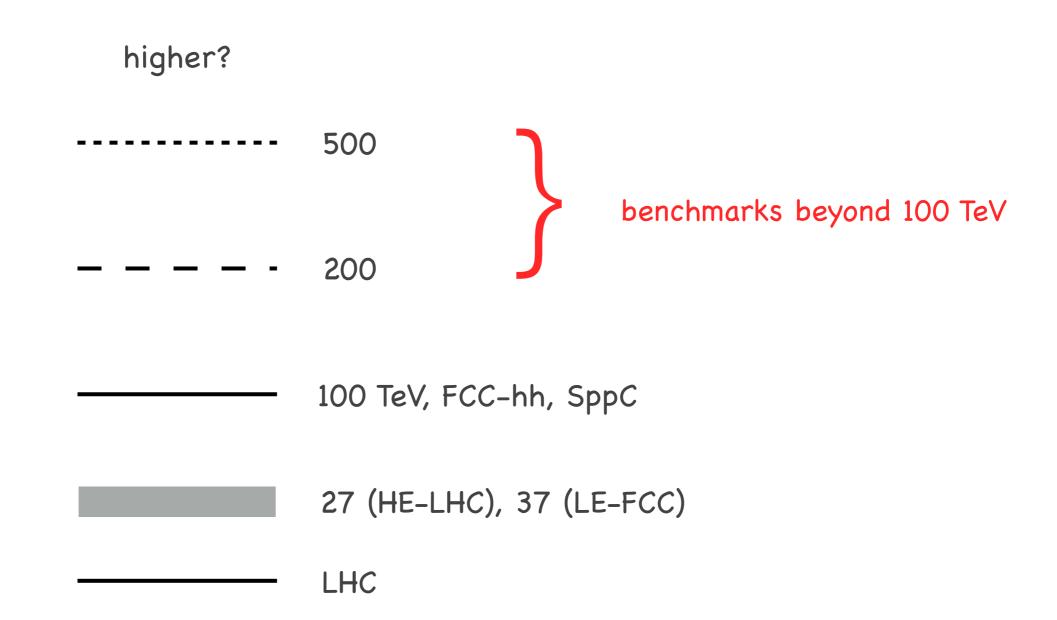
At the same time, the error at hadron colliders typically dominated by systematics, less direct dependence on luminosity.

Based on available studies, 10-30 ab⁻¹ seems to do an adequate job.

Benchmarks beyond known options



Benchmarks beyond known options



Luminosity: 10³⁵–10³⁶ cm⁻² s⁻¹

Example: Naturalness

All Colliders: Top squark projections

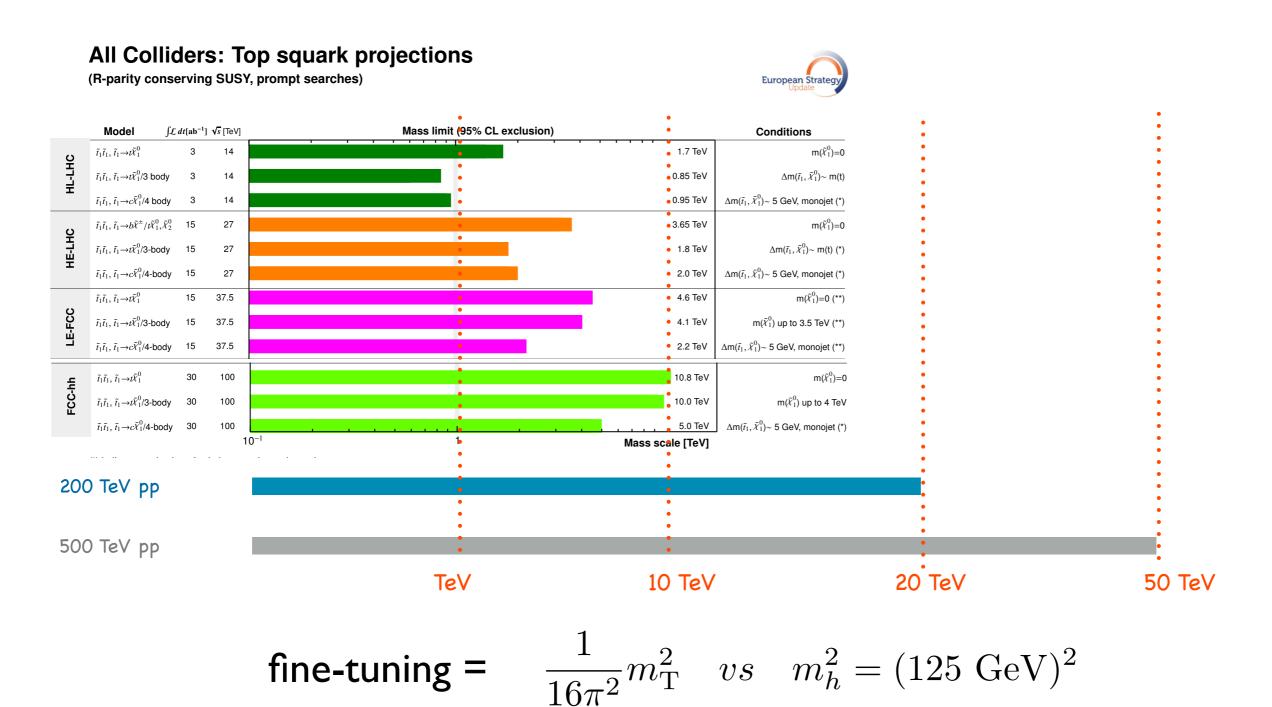
(R-parity conserving SUSY, prompt searches)



	Model ∫⊥	dt[ab ⁻¹]	√s [TeV]	Mass limit (95% CL exclusion)	Conditions
ပ	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 {\rightarrow} t \tilde{\chi}_1^0$	3	14	1.7 TeV	$m(\tilde{\chi}_1^0)=0$
HL-LHC	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 { ightarrow} t \tilde{\chi}_1^0 / 3 { m body}$	3	14	0.85 TeV	$\Delta m(\tilde{t}_1, \tilde{\chi}^0_1) \sim m(t)$
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 { ightarrow} c \tilde{\chi}_1^0 / 4 { m body}$	3	14	0.95 TeV	Δ m $(\tilde{t}_1, \tilde{\chi}_1^0)$ ~ 5 GeV, monojet (*)
ပ	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 {\rightarrow} b \tilde{\chi}^{\pm} / t \tilde{\chi}_1^0, \tilde{\chi}_2^0$	15	27	• 3.65 TeV	$m(\tilde{\chi}_1^0)=0$
HE-LHC	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 {\rightarrow} t \tilde{\chi}_1^0 / 3$ -body	15	27	• 1.8 TeV	$\Delta m(\tilde{t}_1, \tilde{\chi}_1^0) \sim m(t) \ (*)$
I	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 { ightarrow} c \tilde{\chi}_1^0 / 4$ -body	15	27	2.0 TeV	Δ m($\tilde{t}_1, \tilde{\chi}_1^0$)~ 5 GeV, monojet (*)
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 {\rightarrow} t \tilde{\chi}_1^0$	15	37.5	• 4.6 TeV	$m(\tilde{\chi}_1^0)=0 \ (**)$
LE-FCC	$\tilde{t}_1\tilde{t}_1,\tilde{t}_1{ ightarrow}t\tilde{\chi}_1^0/3$ -body	15	37.5	• 4.1 TeV	$m(\tilde{\chi}^0_1)$ up to 3.5 TeV (**)
Ë	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 { ightarrow} c \tilde{\chi}_1^0 / 4$ -body	15	37.5	2.2 TeV	Δ m($\tilde{t}_1, \tilde{\chi}_1^0$)~ 5 GeV, monojet (**)
돈	$\tilde{t}_1\tilde{t}_1,\tilde{t}_1{ ightarrow}t\tilde{\chi}_1^0$	30	100	10.8 TeV	$m(\tilde{\chi}_1^0)=0$
FCC-hh	$\tilde{t}_1\tilde{t}_1,\tilde{t}_1\!\rightarrow\!t\tilde{\chi}_1^0$ /3-body	30	100	10.0 TeV	$m(\tilde{\chi}^0_1)$ up to 4 TeV
	$\tilde{t}_1\tilde{t}_1,\tilde{t}_1{ ightarrow}c\tilde{\chi}_1^0/4 ext{-body}$	30	100	5.0 TeV	$\Delta m(\tilde{t}_1, \tilde{\chi}_1^0) \sim 5 \text{ GeV, monojet (*)}$
				O ⁻¹ Mass scale [TeV]	
				•	
				•	
				•	
				•	
					,
				TeV 10 TeV	

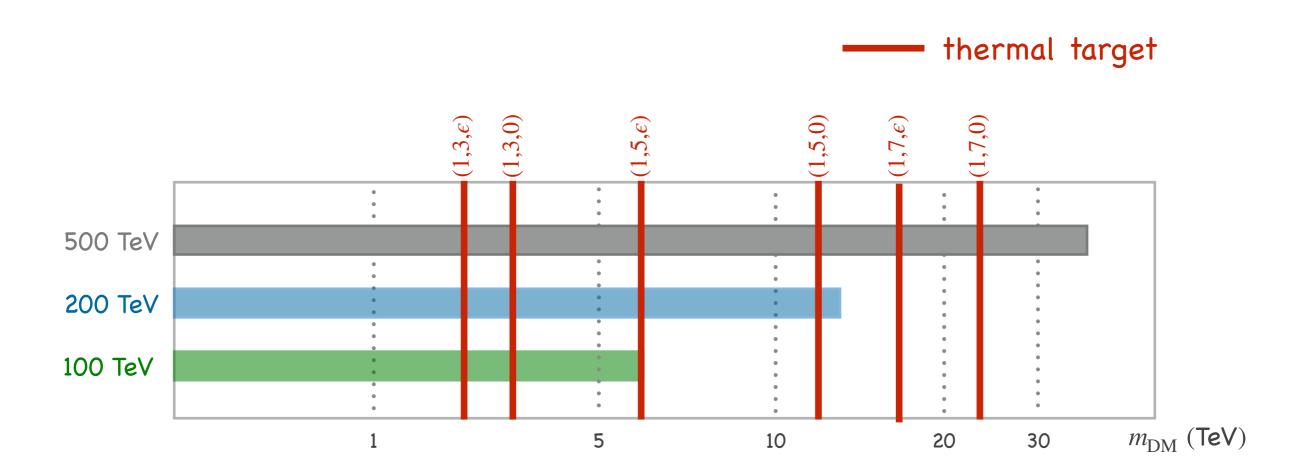
fine-tuning =
$$\frac{1}{16\pi^2}m_{\rm T}^2$$
 vs $m_h^2 = (125~{\rm GeV})^2$

Pushing ahead with hadron collider



Comparing with 100 TeV, gaining a factor of 4(25) at 200(500) TeV pp

EW Dark matter reach



Higher energy needed to cover higher dimensional multiplets.

Either discovery or exclusion, we can make a clear statement of this very compelling WIMP DM scenario.

Hadron vs lepton (intuition)

pp collider

- Higher energy.
- Messier, noisier.
- Probing more interactions.
 Stronger if NP has strong interaction.

lepton collider

- Lower energy.
- Cleaner environment,
 better sensitivity,
 precision.
- Stronger for electroweak states.

Hadron vs lepton (intuition)

pp collider

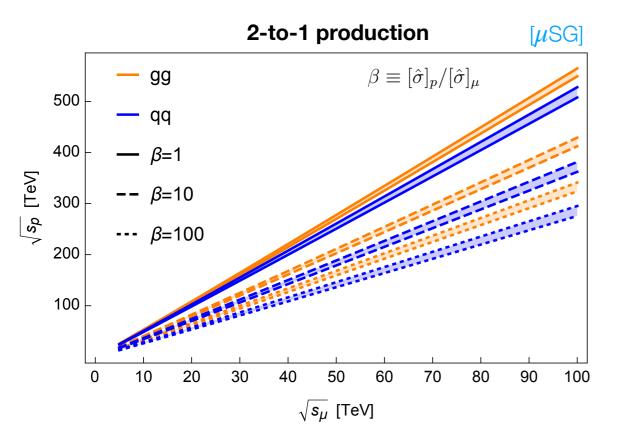
- Higher energy.
- Messier, noisier.
- Probing more interactions.
 Stronger if NP has strong interaction.

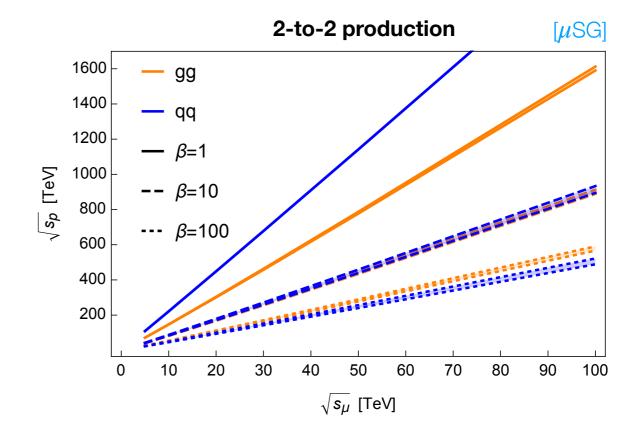
lepton collider

- Lower energy.
- Cleaner environment,
 better sensitivity,
 precision.
- Stronger for electroweak states.

However, this comparison really depends on what is achievable.

Hadron vs lepton



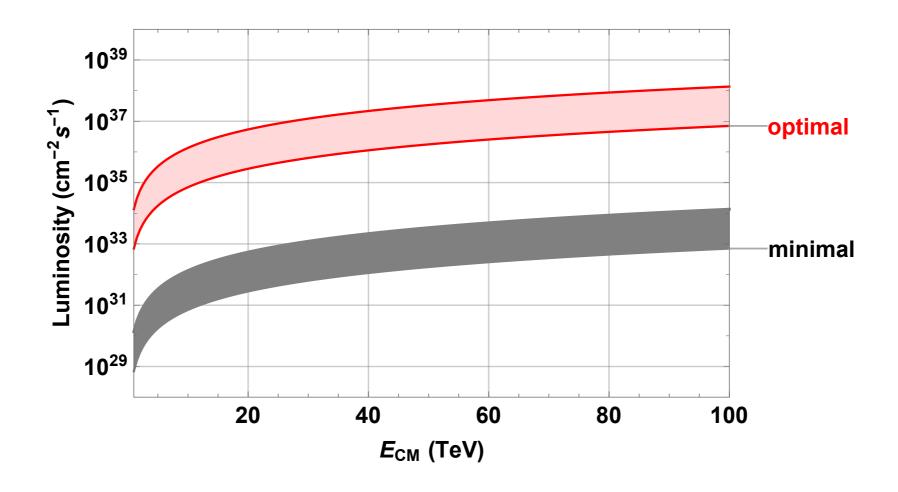


We "know" how to make hadron colliders.

But, we also need it to be at much (O(10)) higher E_{CM} .

What is our best route to (super) high energies?

Conclusion: Lepton collider



Beyond the proposed Higgs factories, high energy lepton colliders from TeV to 10s TeV extremely interesting!

A bare minimal scenario with much less lumi can already do some interesting searches.

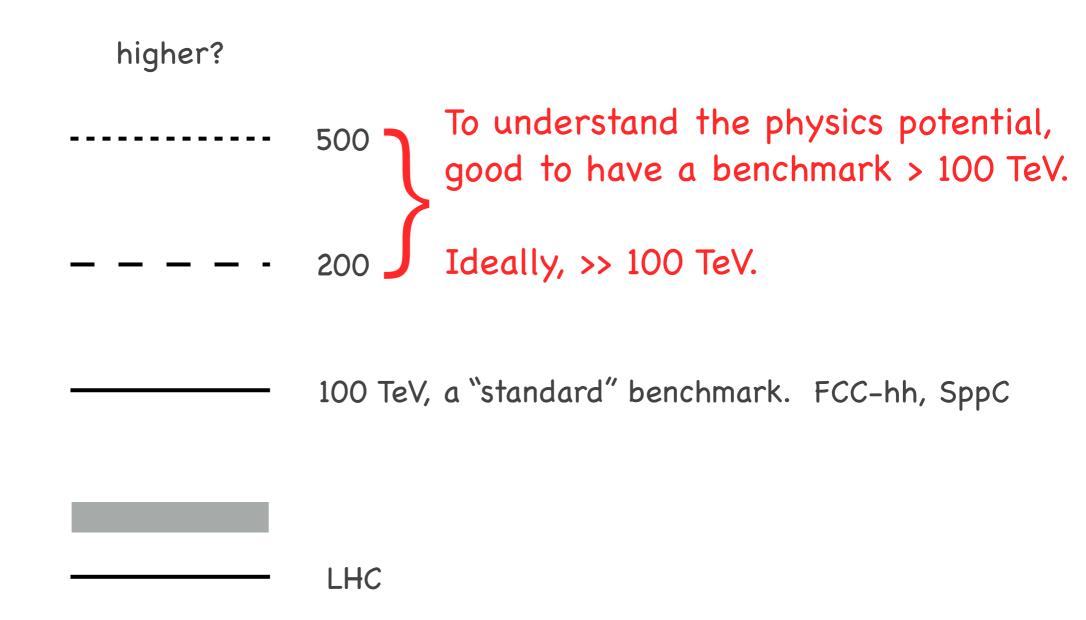
Conclusion: Lepton collider

Luminosity cm ⁻² s ⁻¹	I.5 TeV	3 TeV	6 TeV	10 TeV	I4 TeV	30 TeV	100 TeV
Direct search minimal	3×10 ²⁹	1030	5×10 ³⁰	2×10 ³¹	5×10 ³¹	2×10 ³²	2×10 ³³
Direct search optimal	3×10 ³³	1034	5×10 ³⁴	2×10 ³⁵	5×10 ³⁵	2×10 ³⁶	2×10 ³⁷
Precision minimal	3×10 ³⁰	8x10 ³¹	2×10 ³³	1034	5×10 ³⁴	1036	2×10 ³⁷
Precision optimal	7×10 ³²	I 0 ³⁴	2×10 ³⁵	2×10 ³⁶	5×10 ³⁶	1038	2×10 ³⁹

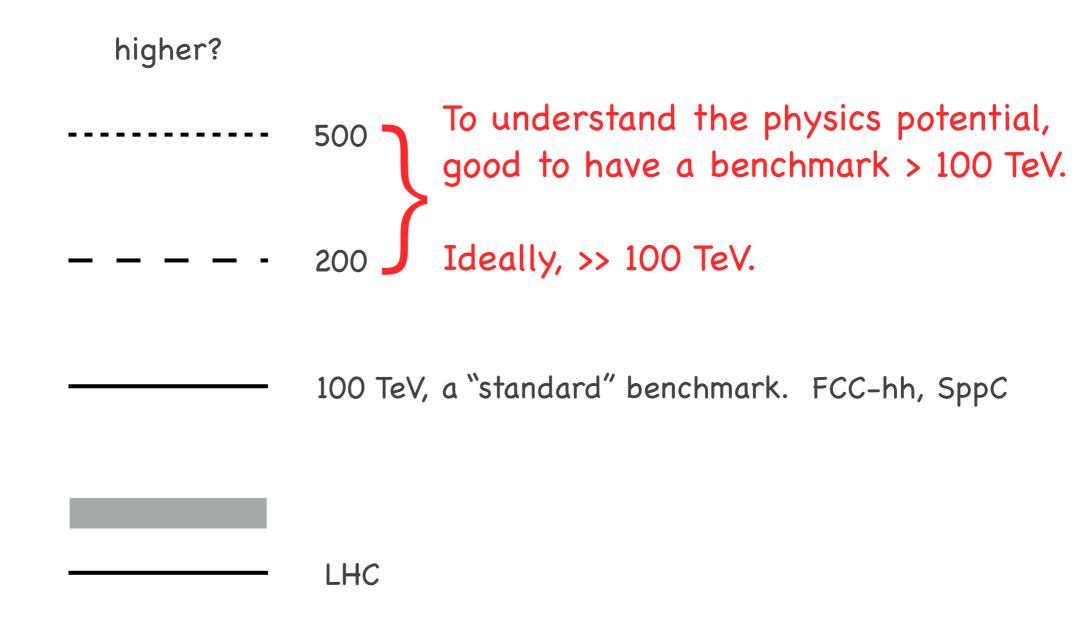
A bare minimal scenario with much less lumi can already do some interesting searches.

Upgrades towards optimal scenarios highly desired, covering interesting cases such as dark matter, and allow precision measurements to reach its full potential.

Conclusion: hadron collider



Conclusion: hadron collider



Luminosity at HL-LHC level can already have interesting sensitivity.

Higher luminosity at 10^{35} – 10^{36} cm⁻² s⁻¹ can realize full physics potential.

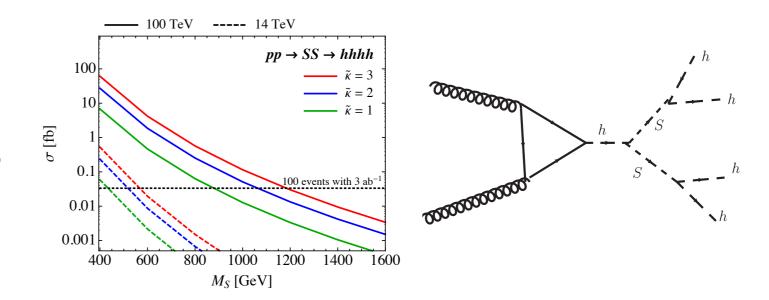
extras

Probing neutral naturalness

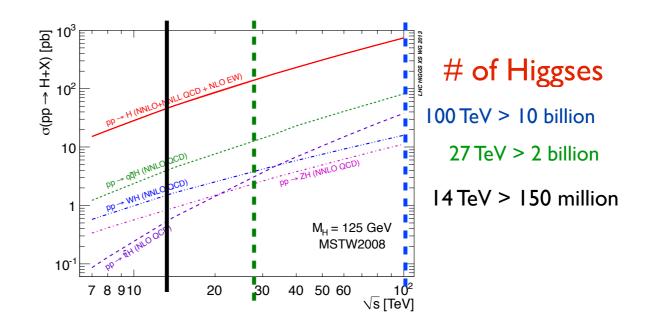
100 TeV Testing tuning to 10%

200(500) push down to several %

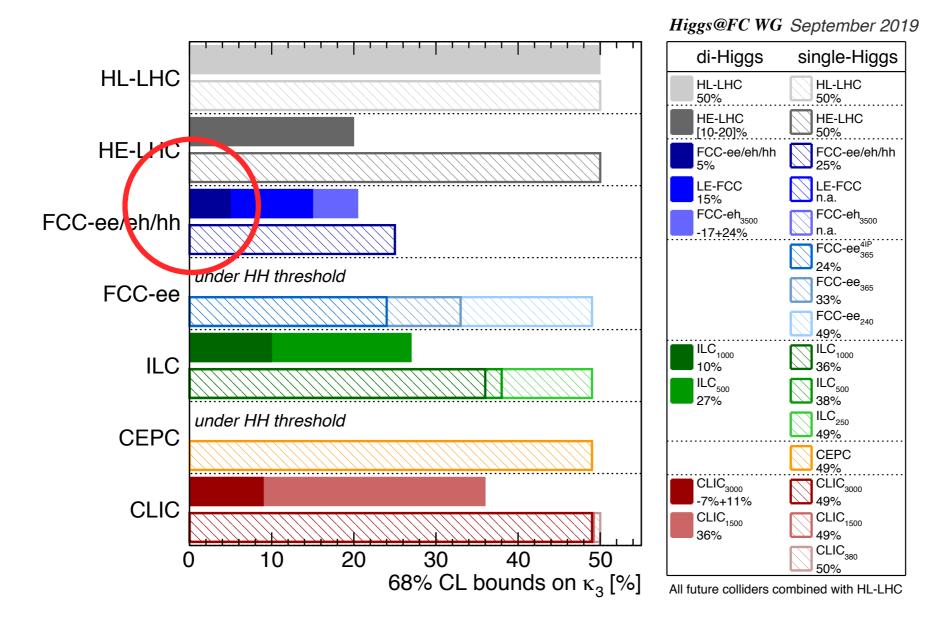
Probing m_s > TeV



100 TeV collider can improve at least a factor of 10 beyond the LHC.

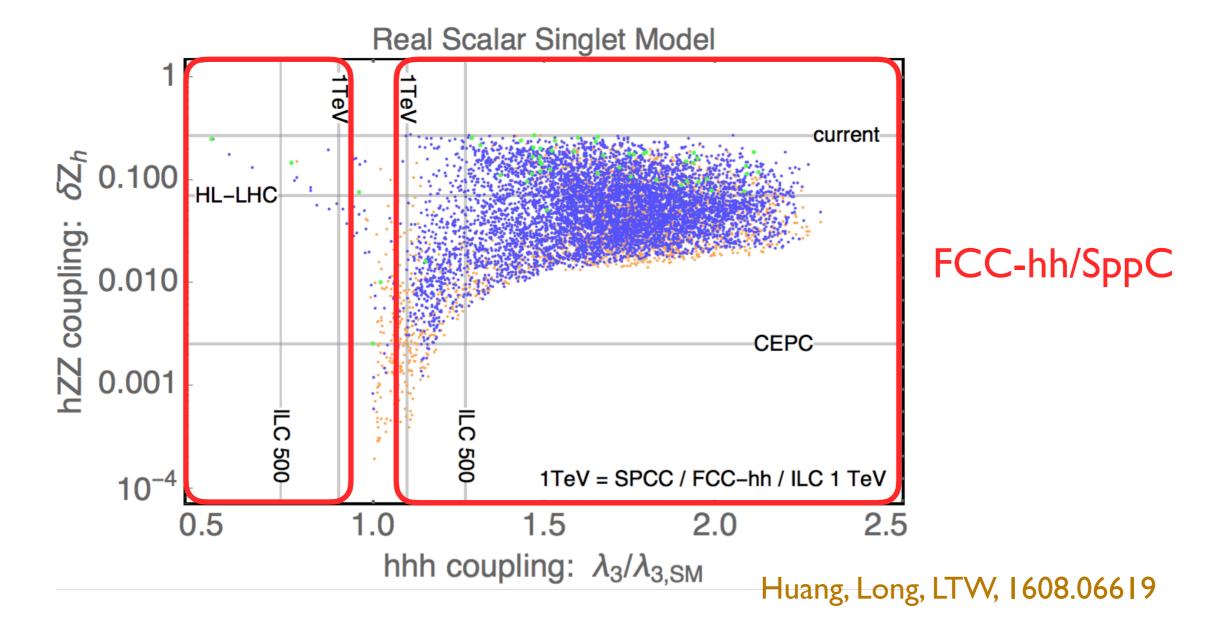


Higgs potential



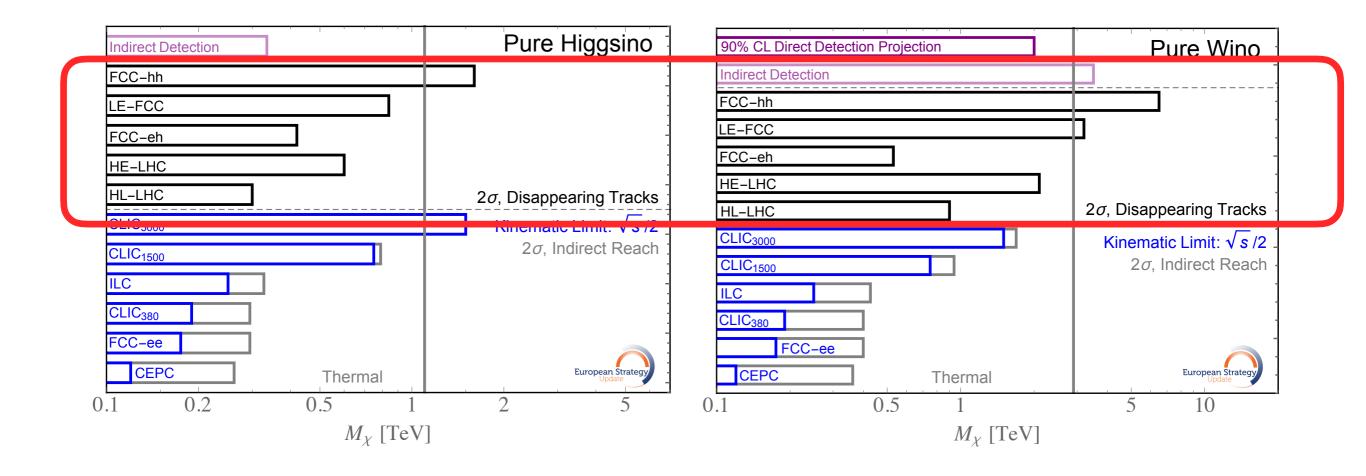
Extrapolating to higher energies more difficult. We should expect a factor of a few improvement.

Probing EW phase transition



Orange = first order phase transition, $v(T_c)/T_c > 0$ Blue = "strongly" first order phase transition, $v(T_c)/T_c > 1.3$ Green = very strongly 1PT, could detect GWs at eLISA

Dark matter reach



100 TeV pp collider is needed to cover the EW doublet (Higgsino) and triplet (wino) DM.

Some Future Hadron collider proposals



parameter	FCC-hh		FCC- hh-6T	HE-LHC	HL- LHC	LHC
collision energy cms [TeV]	1	100	37.5	27	14	14
dipole field [T]	16		6	16	8.33	8.33
beam current [A]		0.5	0.6	1.1	1.1	0.58
synchr. rad. power/ring [kW]	2400		57	101	7.3	3.6
peak luminosity [10 ³⁴ cm ⁻² s ⁻¹]	5 30		10 (lev.)	16	5 (lev.)	1
events/bunch crossing	170 1000		~300	460	132	27
stored energy/beam [GJ]		8.4	3.75	1.4	0.7	0.36

- NbTi technology from LHC, magnet with single-layer coil providing 6 T at 1.9 K:
 - → Corresponding beam energy 18.75 TeV or 37.5 TeV c.m.
 - → Significant reduction of synchrotron radiation wrt FCC-hh (factor 50) and corresponding cryogenic system requirements.
- Luminosity goal 10 ab⁻¹ over 20 years or 0.5 ab⁻¹ annual luminosity:
 - → Beam current 0.6 A or 20% higher than for FCC-hh, 1.2E11 ppb (FCC-hh: 1.0 ppb).
 - → Stored beam energy 3.75 GJ vs 8.4 GJ for FCC-hh.
- Analysis of physics potential, technology requirements and cost ongoing.

M. Benedikt and F. Zimmermann, FCC week

Future Hadron colliders

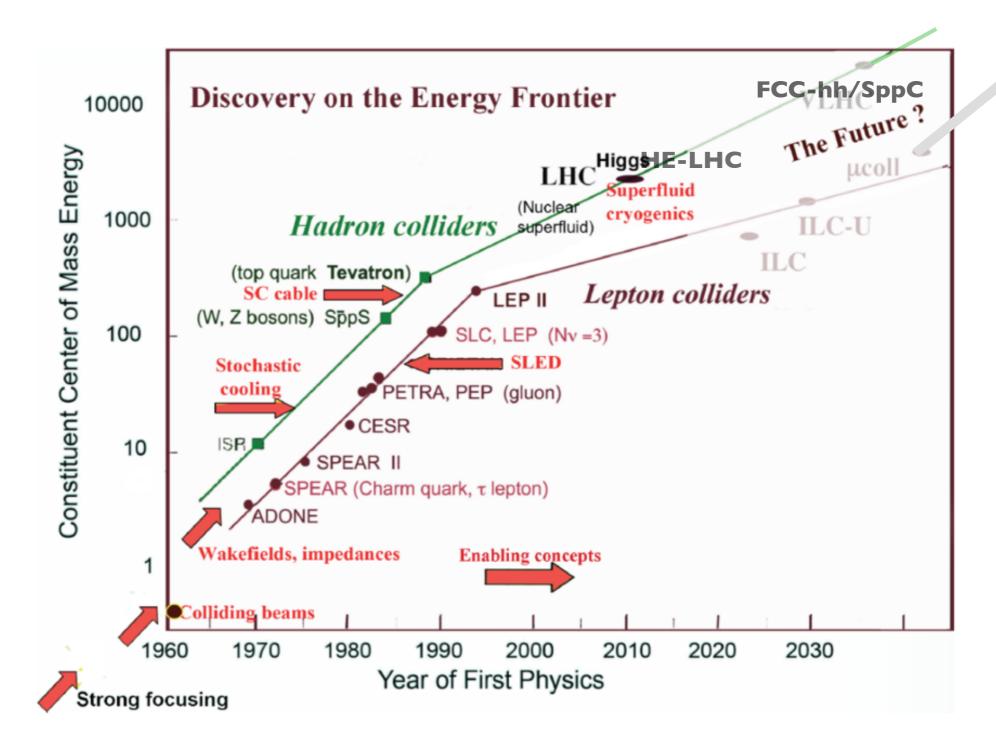
Hadron collider parameters (pp)

parameter	F	CC-hh	HE-LHC	(HL) LHC	
collision energy cms [TeV]		100	27	14	
dipole field [T]		16 16		8.3	
circumference [km]		100	27	27	
beam current [A]		0.5	1.12	(1.12) 0.58	
bunch intensity [10 ¹¹]	•	1 (0.5)	2.2	(2.2) 1.15	
bunch spacing [ns]	25 (12.5)		25 (12.5)	25	
norm. emittance $\gamma ε_{x,y}$ [μm]	2.2 (1.1)		2.5 (1.25)	(2.5) 3.75	
IP β* _{x,y} [m]	1.1 0.3		0.25	(0.15) 0.55	
Iuminosity/IP [10 ³⁴ cm ⁻² s ⁻¹]	5 30		28	(5) 1	
peak #events / bunch Xing	170 1000 (500)		800 (400)	(135) 27	
stored energy / beam [GJ]	8.4		1.4	(0.7) 0.36	
SR power / beam [kW]	2400		100	(7.3) 3.6	
transv. emit. damping time [h]	1.1		3.6	25.8	
initial proton burn off time [h]	17.0	3.4	3.0	(15) 40	

target luminosity HL-LHC: 3 ab-1, HE-LHC and FCC-hh: 20-30 ab-1

Collider	Type	\sqrt{s}	\$\mathcal{P}[\%]	N_{Det}	$\mathscr{L}_{ ext{inst}}$ /Det.	\mathscr{L}	Time	Ref.
			$[e^{-}/e^{+}]$		$[10^{34} \text{cm}^{-2} \text{s}^{-1}]$	$\left[ab^{-1} \right]$	[years]	
HL-LHC	pp	14 TeV	_	2	5	6.0	12	[23]
HE-LHC	pp	27 TeV	_	2	16	15.0	20	[23]
FCC-hh	pp	100 TeV	_	2	30	30.0	25	[637]
FCC-ee	ee	M_Z	0/0	2	100/200	150	4	[637]
		$2M_W$	0/0	2	25	10	1-2	
		240 GeV	0/0	2	7	5	3	
		$2m_{top}$	0/0	2	0.8/1.4	1.5	5	
	(1y SD before	$re 2m_{top} run$)			(+1)	
ILC	ee	250 GeV	±80/±30	1	1.35/2.7	2.0	11.5	[342]
		350 GeV	$\pm 80/\pm 30$	1	1.6	0.2	1	[346]
		500 GeV	$\pm 80/\pm 30$	1	1.8/3.6	4.0	8.5	
	(1	y SD after 2	250 GeV rur	n)			(+1)	
CEPC	ee	M_Z	0/0	2	17/32	16	2	[509]
		$2M_W$	0/0	2	10	2.6	1	
		240 GeV	0/0	2	3	5.6	7	
CLIC	ee	380 GeV	±80/0	1	1.5	1.0	8	[638]
		1.5 TeV	$\pm 80/0$	1	3.7	2.5	7	
		3.0 TeV	$\pm 80/0$	1	6.0	5.0	8	
	(2y 3	SDs betwee	n energy sta	ges)			(+4)	
LHeC	ep	1.3 TeV	_	1	0.8	1.0	15	[636]
HE-LHeC	ep	1.8 TeV	_	1	1.5	2.0	20	[637]
FCC-eh	ep	3.5 TeV	_	1	1.5	2.0	25	[637]

Why hadron collider?



Highest energies achieved in the lab. Offers a first direct glance at shortest distances.